A model of the X-ray response of the ACIS CCD

G. Prigozhin[a], A. Rasmussen[b], M. Bautz[a], G. Ricker[a]

[a]Massachusetts Institute of Technology, Center for Space Research, Cambridge, MA

[b]Columbia University, Columbia Astrophysics Laboratory, New York

ABSTRACT

We have developed a comprehensive model of the response of a CCD to soft X-ray illumination. The model is based on the Monte Carlo technique and follows the interactions with the device material of individual photons thrown into the structure, calculating device reaction to each of them. It incorporates a very detailed description of the CCD gate structure, as well as accurately measured absorption coefficients. The fluorescent and escape peak model takes into account interactions inside the gate structure, which dramatically improves the agreement with the experimental data at energies close to the Si absorption edge. The shape of the low energy tail is simulated according to our new model of electron cloud charge splitting at the interface between Si and SiO_2. An origin of the tail in the horizontally split events is explained as coming from the $p+$ area in the channel stop region and is modeled accordingly.

Keywords: charge coupled devices, response function, X-rays, quantum efficiency

1. INTRODUCTION

The focal plane of the AXAF Imaging Spectrometer (ACIS) is populated with 10 Charge Coupled Devices (CCD) which convert incoming X-ray photons into electric signal and allow to get images of the X-ray sources simultaniously supplying information about the spectral composition of the input flux. The spectral structure of the celestial sources is very complicated and in order to succesfully deconvolve the source spectrum from the output signal of a CCD it is very important to have a high fidelity model of the device response.

We have developed a detailed model of the frontside illuminated CCD response to the X-ray illumination. The model is based on the most common approach for multipixel detectors sumulation (see, for instance,[1]) in which the photons of given energy are thrown into the random positions on the CCD surface and the device reaction is calculated for each individual photon. The procedure is then repeated many times to produce a histogram of the device response. Some blocks of our model, such as electron cloud diffusion in both depleted and undepleted bulk, charge cloud splitting between the pixels, event finding procedures, are similar to the ones described before in[1–3] and we will not discuss them in detail. In this work we will focus on the new developments in the device physics which allowed us to produce a model that is more adequate to the stringent calibration requirements of AXAF. The most important new features of our model are discussed below.

2. ABSORPTION IN THE GATE STRUCTURE

The gate structure of the ACIS CCD consists of the thin layers of polysilicon, silicon dioxide, and silicon nitride. X-ray photon can be absorbed in any of these layers with an exponentially distributed (depthwise) probability of interaction. To model this process the simulation program generates an exponentially distributed random number R. If $R > \frac{d}{\lambda}$ we assume that the photon have passed the layer of thicknness d with characteristic absorption length λ without being absorbed. Otherwise the photon is absorbed inside the layer.

The cross-section of the gate structure of the ACIS CCD along the direction of charge transfer is shown schematically on Fig. 1. It shows the details of the gate structure which were measured experimentally with Scanning Electron Microscope (SEM). Another cross-section, perpendicular to the direction of the charge transfer is shown on Fig. 2.

Absorption of X-rays in the gates becomes a very significant factor at low energies (below 1 keV) and immediately above silicon edge. Our model simulates all the details of the gate structure shown in Figures 1,2, such as gate overlaps,

Corresponding author G.P.; e-mail: gyp@space.mit.edu