

HETG/LETG — Status

Chandra Quarterly Review No. 57

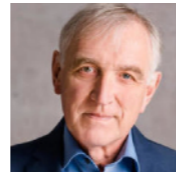
25 June 2024

David Huenemoerder
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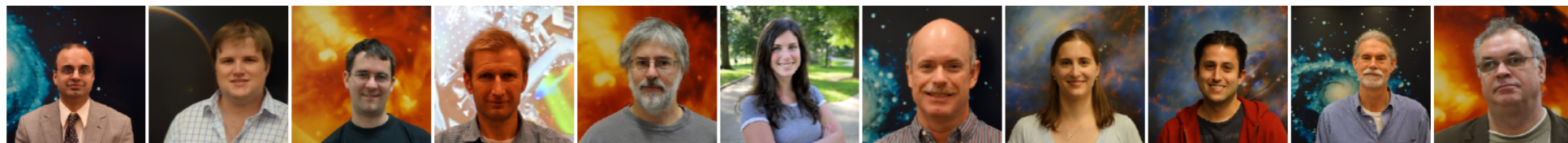
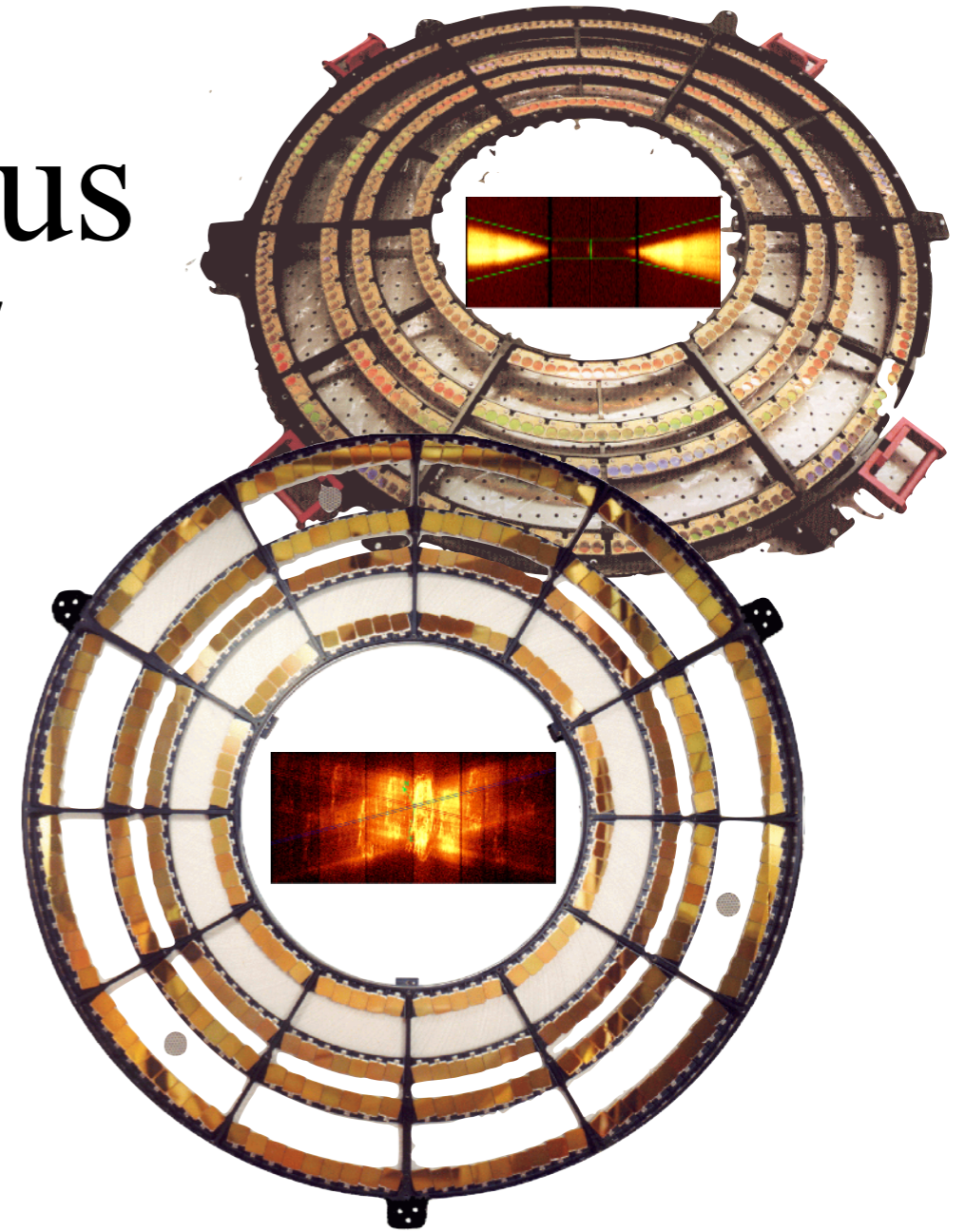
HETG IPI: Prof. Claude R. Canizares
MIT Kavli Institute



LETG IPIs: Dr. Peter Predehl
Max Planck Institute



Dr. Jelle Kaastra
SRON



faculty, staff;



postdocs

Proposal Cycle 24 Jan 2023 (341 / 474 ks)

- ★ XRB: 4U 1624-490 29/135 ks Accretion disk structure (with NuSTAR, XRISM)
- ★ XRB: Cen X-3 48 ks Eclipsing X-ray pulsar; accretion (ongoing - low visibility)
- ★ AGN: MCG-6-30-15 224 ks Time-dependent photoionisation modeling of outflows
- ★ BH: SS 433 39/60 ks Relativistic jet physics (coordinated with HRC, Swift GO)

Proposal Cycle 25 Jan 2024 (82 / 758 ks) (total of 1208 ks with all TOO)

- ★ ULX/NS: M33 X-8 0/85 ks Pulsar wind outflow, absorption lines
- ★ Stars: ζ Puppis 0/190 ks Stellar winds, long-term monitoring
- ★ XRB: 4U 1626-67 0/94 ks Ultra-compact binary pulsar line shapes and torque-reversal
- ★ XRB: Her X-1 0/50 ks Pulsar accretion disk/wind dynamics during outburst (TOO)
- ★ ISM: GX 9+9 82/190 ks Galactic silicon absorption edge survey
- ★ BH: GRS 1915+105 0/100 ks Micro-quasar (TOO; coordinated with NICER, NuSTAR)
- ★ GW Transient: 0/300 ks Gravitational Wave source followup (TOO)
- ★ NS: Terzan 5 X-2 0/200 ks Neutron Star outburst (TOO) (a.k.a. IGR J17480-2446)

Jargon list:
 AGN: Active Galactic Nucleus
 BH: Black Hole
 ISM: InterStellar Medium
 NS: Neutron Star
 SN: SuperNova
 ULX: Ultra-Luminous X-ray source
 SNR: SuperNova Remnant
 XRB: X-ray Binary
 LMXB: Low Mass XRB



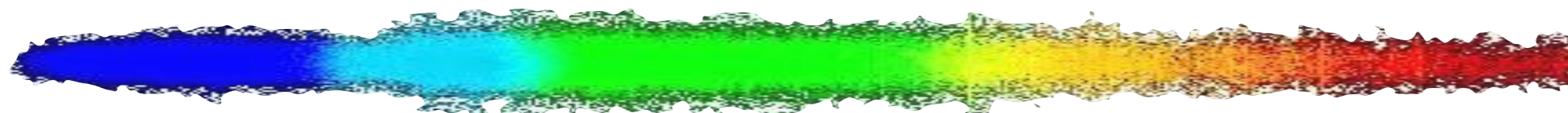
Proposal Cycle 24: start Jan 2023 (146 / 338 ks)

★ AGN: (Predehl/MPE)	WISEA J050222.16-341201.6	25 ks	Luminous $z > 5.6$ quasars from eROSITA (ACIS-S)
★ AGN: (Predehl/MPE)	WISEA J230341.02-542730.6	0/32	Luminous $z > 5.6$ quasars from eROSITA (ACIS-S)
★ AGN: (Predehl/MPE)	WISEA J050411.92-254959.0	26	Luminous $z > 5.6$ quasars from eROSITA (ACIS-S)
★ Galaxies: (Predehl/MPE)	eFEDSJ083933	84	Shocks in an eROSITA detected galaxy cluster (ACIS-S)
★ AGN: (Kaastra/SRON)	NGC 3783	10/166	Outflows, variability (with XRISM) (ACIS-S/HETG)

Proposal Cycle 25 start Jan 2024 (115 / 343 ks)

★ ISM: (Predehl/MPE)	HD 115247	57 ks	Galactic bubbles absorption (LETG/HRC-S; w/ eROSITA)
★ ISM: (Predehl/MPE)	HE1338-1423	58 ks	Galactic bubbles absorption (LETG/HRC-S; w/ eROSITA)
★ ISM: (Predehl/MPE)	LEDA407	0/59 ks	Galactic bubbles absorption (LETG/HRC-S; w/ eROSITA)
★ AGN: (Kaastra/SRON)	A1550	0/166 ks	Outflows, variability

AGN: Active Galactic Nucleus
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Performance Nov 2023 — May 2023

HETG/ACIS-S: 845 ks

- 46 observations on 13 targets (17 GO, 19 GTO, 8 DDT, 2 Cal)

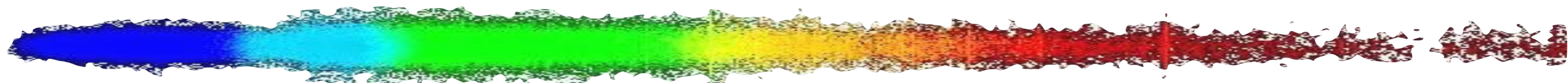
LETG:

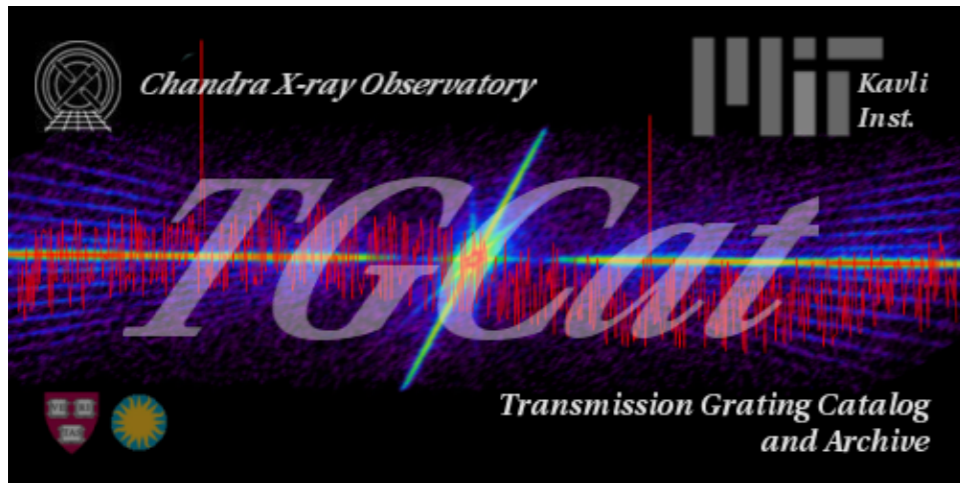
- 16 LETG/HRC-S observations, 4 targets (189 ks; 4 GO, 10 GTO, 2 Cal)
- 12 LETG/ACIS-S observations, 1 target (351 ks; 12 Cal — contamination)
- 1 LETG/HRC-I observation, 1 target (2 ks; 1 Cal)

All instruments: 11.1 Ms

Grating instruments: 1.4 Ms (13%)

Grating performance is nominal.

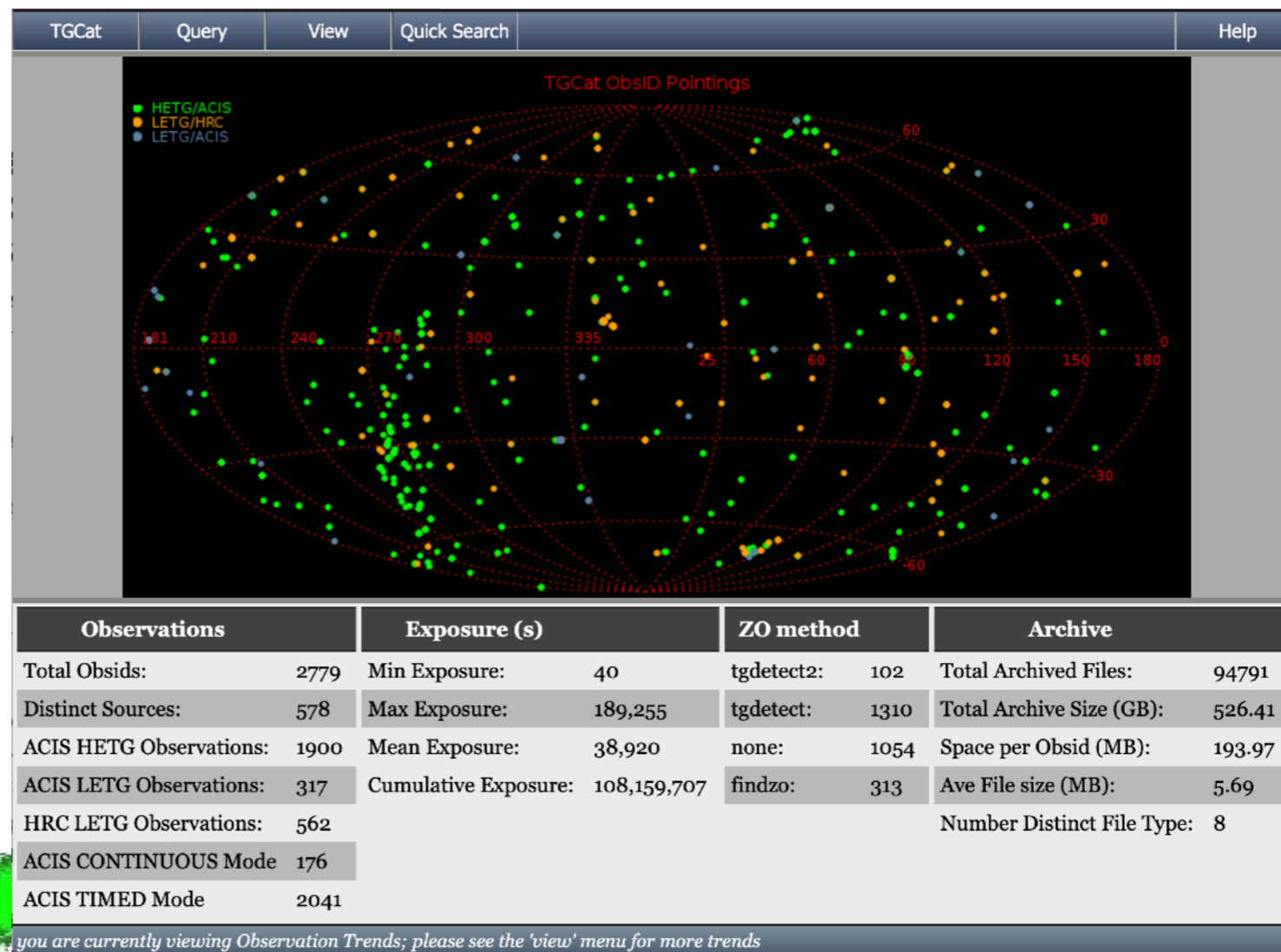




<http://tgcat.mit.edu>

Security updates / migration finished.
Back on-line as of December, 2023.

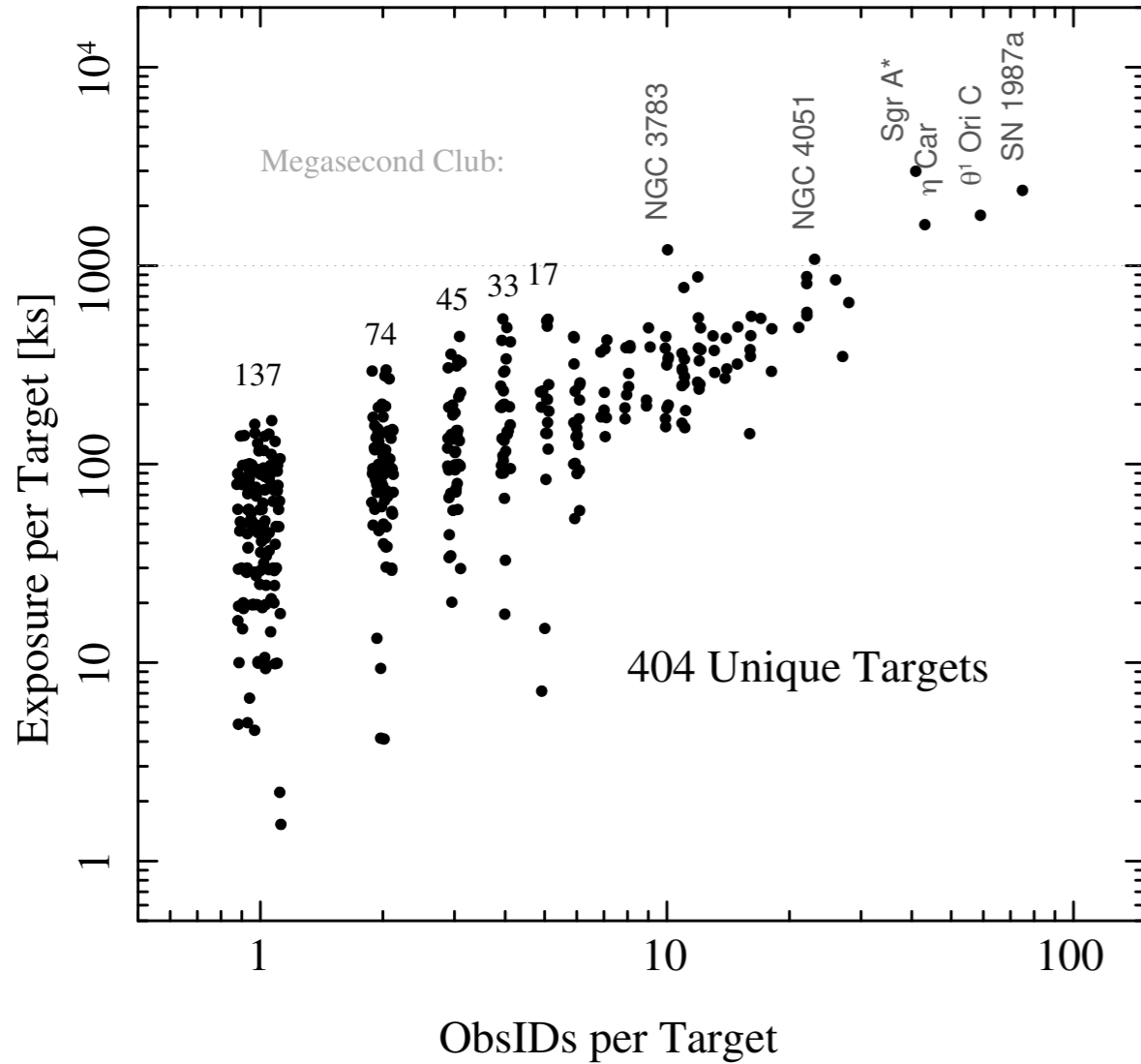
<<http://tgcat.mit.edu/dev/tgcat/tgTrend.php>>



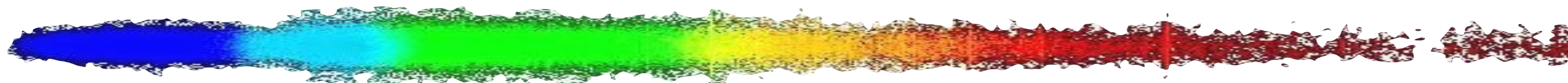
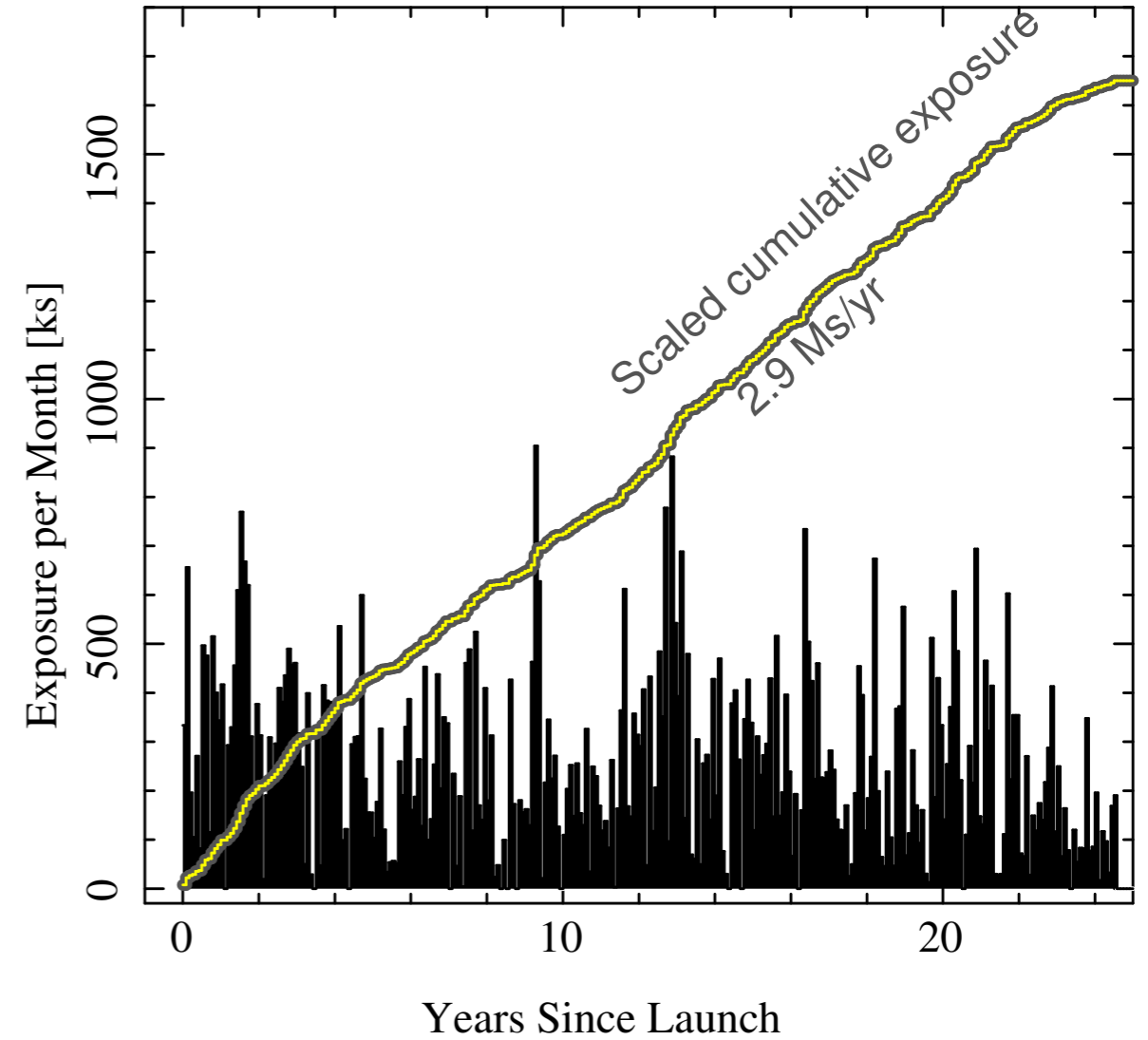
HETG Observations Historical Trends



Exposure per target, and number of visits



Exposure per month, and cumulative exposure



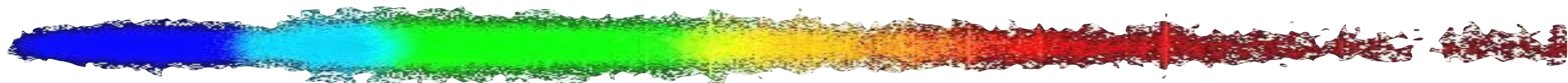
ACIS High-T Operations: OK to use HETG

(but for warmest temperatures, offset to low CHIPY).

Warning in Verification & Validation report if limits exceeded.

HETG 2nd and 3rd order efficiencies: in progress.

Line-spread-function parameters for HETG/HRC-I and off-axis pointings: in progress.





OPEN ACCESS



Chandra HETG X-Ray Spectra and Variability of π Aqr, a γ Cas-type Be Star

David P. Huenemoerder¹ , Pragati Pradhan² , Claude R. Canizares¹ , Sean Gunderson¹ , Richard Ignace³ , Joy S. Nichols⁴ , A. M. T. Pollock⁵ , Norbert S. Schulz¹, Dustin K. Swarm⁶ , and José M. Torrejón⁷

HETGS 100ks
Cycle 23
(6 Observations,
Aug-Oct 2022)

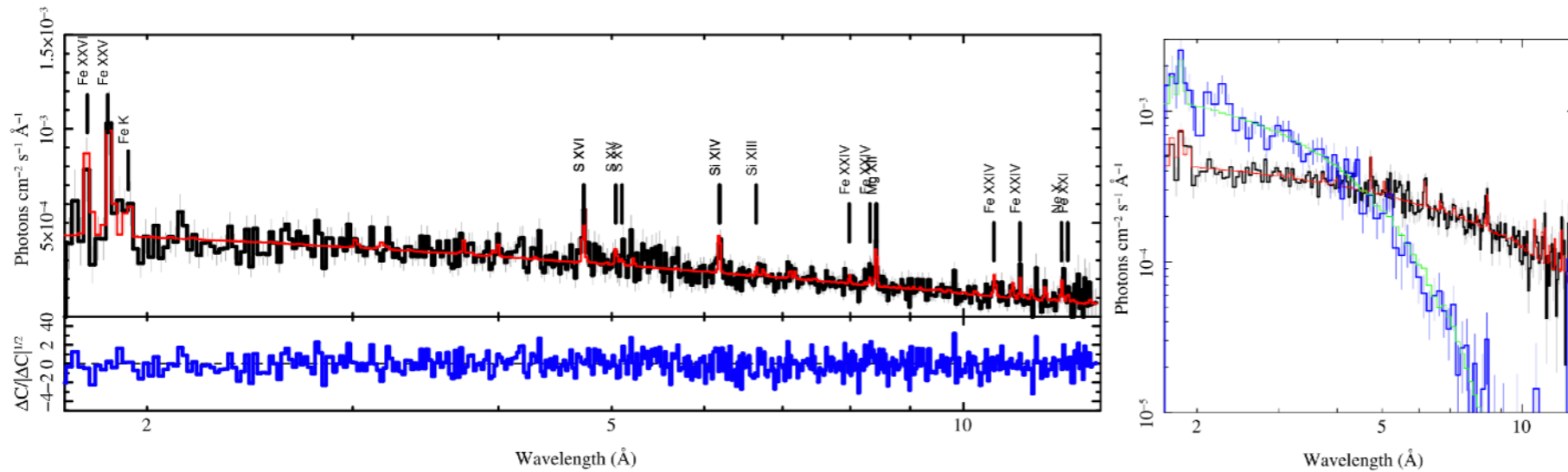
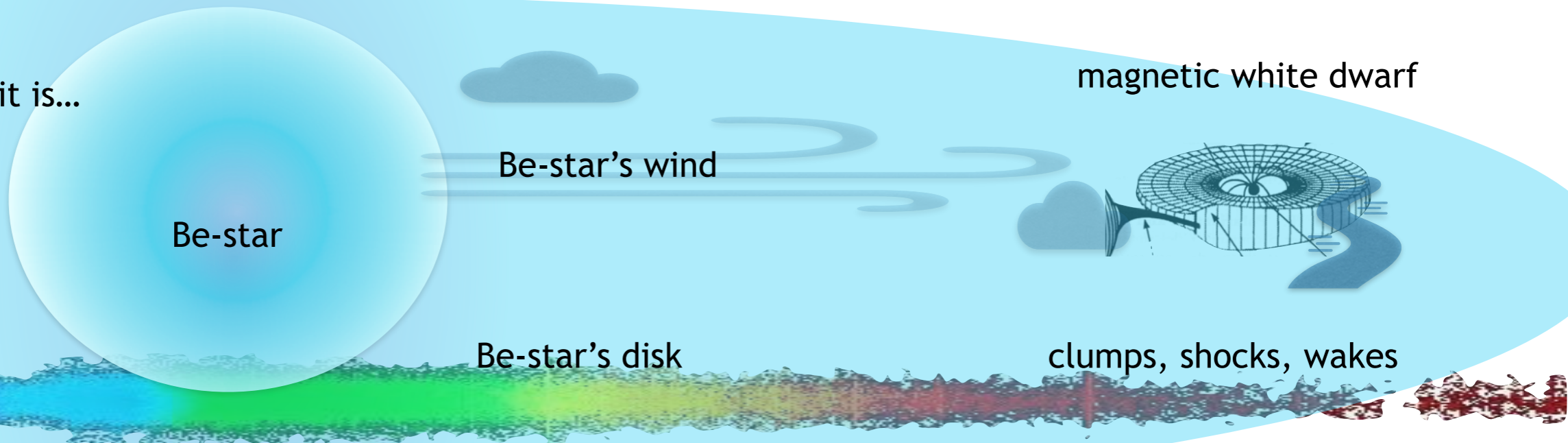


Figure 1. Left: HETG spectrum of π Aqr in black, which is the merged HEG and MEG first orders, excluding the high-hardness ratio observation (27325). In red is a model using three APEC components plus a Gaussian for the putative Fe K fluorescence. On the right, we compare the high-hardness ratio state spectrum (blue) to the rest (black and red).

Mysterious γ Cas type stars: Optically a Be star. In X-rays, MUCH more luminous than typical B-stars, and VERY hot.

What we think it is...



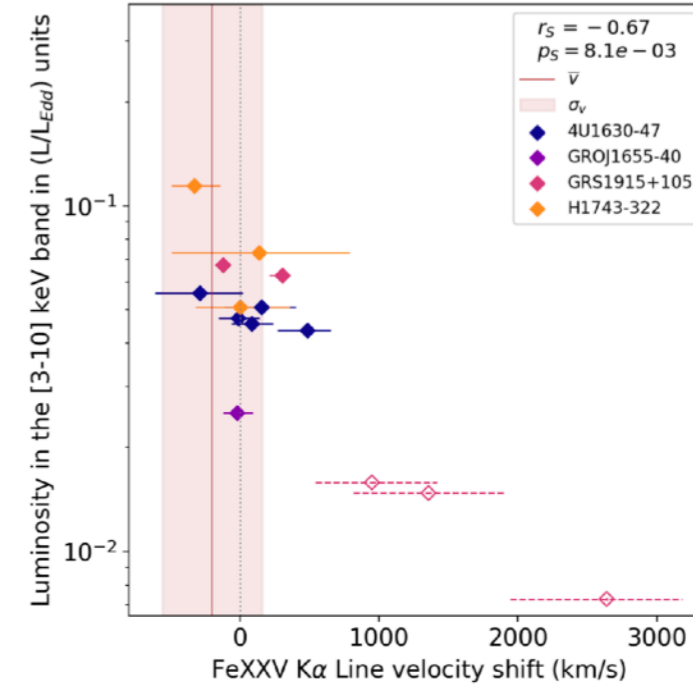
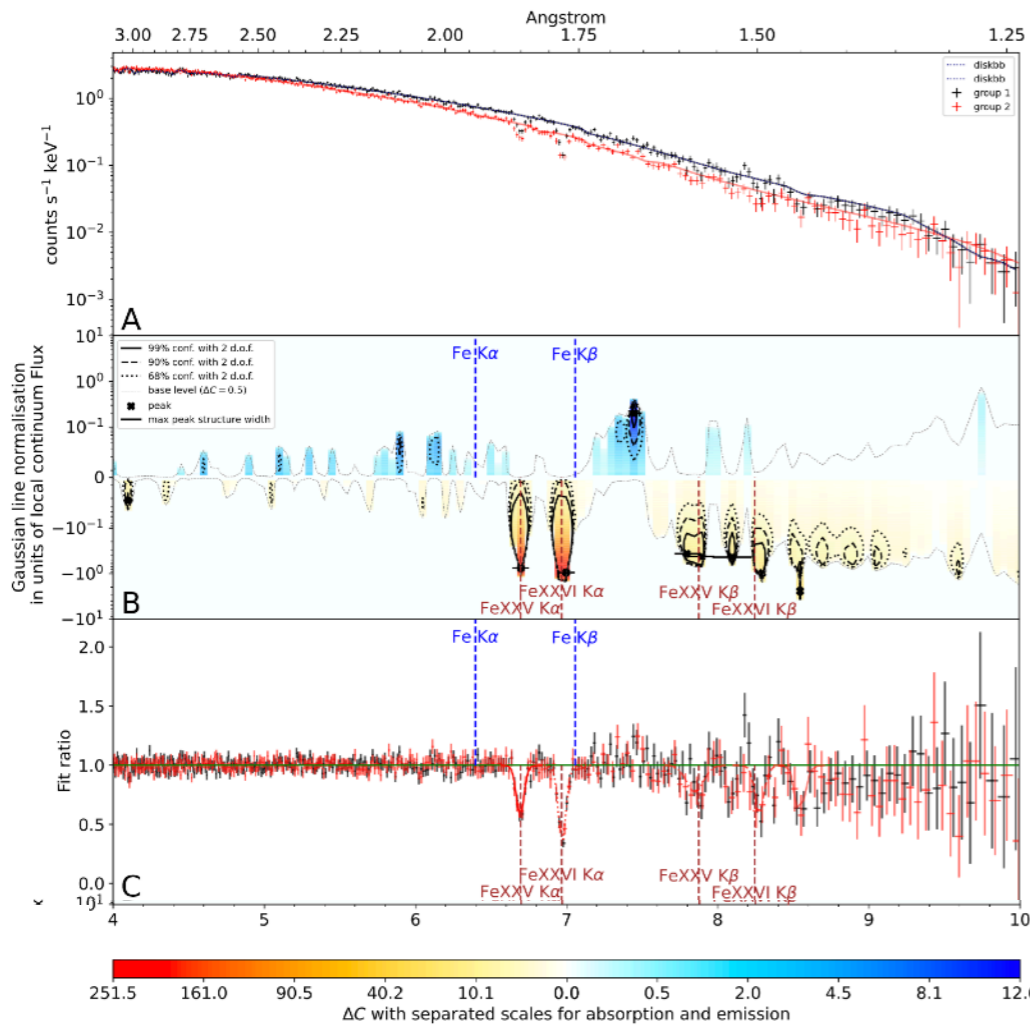
Some Scientific Results Published in the Past 6 Months (2/2)



The current state of disk wind observations in BHLMXBs through X-ray absorption lines in the iron band

M. Parra^{1,2}, P.-O. Petrucci¹, S. Bianchi², V. E. Gianolli^{1,2}, F. Ursini², and G. Ponti^{3,4}

(For about 40 sources reprocessed from XMM archives, or downloaded from TGCat.)



“With the *Chandra* instrument, which proves to be the only instrument sufficiently precise to reliably measure the outflow velocity, the absorption signatures show a global trend of very small blueshifts.”



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