



MIT Kavli Institute

Chandra X-Ray Center

MEMORANDUM

February 2, 2017

To: Jonathan McDowell, SDS Group Leader

From: Glenn E. Allen, SDS

Subject: Temperature-dependent RMF spec

Revision: 1.00

URL: http://space.mit.edu/CXC/docs/docs.html#trmf

File: /nfs/inconceivable/d0/sds/specs/mkacisrmf/trmf_spec_1.00.tex

1 Description

This spec is a high-level description of how a script could be used to create a temperature-dependent RMF.

2 Parameters

The parameters for the script are identical to the parameters for the tool mkacisrmf, except the script includes the two additional parameters:

- 1. gtifile,f,a," ",,,"Input GTI file (e.g. the evt2 file)"
- 2. mtlfile,f,a,"",,,"Input mission timeline file"

3 Processing

- 1. The start and stop times for each one of the $N_{\rm gti}$ good-time intervals for the CCD_ID of interest are read from the appropriate GTI HDU of the gtifile. Hereafter these times for the $i^{\rm th}$ interval are referred to as START[i] and STOP[i], respectively.
- 2. The $N_{\rm mtl}$ sets of values for the time and focal-plane temperature are read from the mtlfile. Hereafter, the $j^{\rm th}$ time and temperature are referred to as ${\tt TIME_{mtl}}[j]$ and ${\tt FP_TEMP_{mtl}}[j]$, respectively. The header of the mtlfile is read to obtain the values of the keywords ${\tt TIMEDEL}$ and ${\tt TIMEPIXR}$, which are hereafter referred to as ${\tt TIMEDEL_{mtl}}$ and ${\tt TIMEPIXR_{mtl}}$, respectively.
- 3. The CALDB is queried to obtain the N_{resp} response files associated with the time of the observation. There is one file for each temperature range calibrated. Hereafter, the k^{th} response file is referred to as respfile[k].

- 4. The minimum and maximum focal-plane temperatures for each respfile are obtained from header keywords in the file. Hereafter these temperatures for the k^{th} file are referred to as, FP_TEMP_{min}[k] and FP_TEMP_{max}[k], respectively.
- 5. For each good-time interval $i = 0, 1, ..., N_{\text{gti}} 1$:
 - (a) For each focal-plane temperature $j=0,1,\ldots,N_{\rm mtl}-1$:

i.

$$t_0 = \text{TIME}_{\text{mtl}}[j] - \text{TIMEPIXR}_{\text{mtl}} \times \text{TIMEDEL}_{\text{mtl}}.$$
 (1)

ii. If

$$t_0 < START[i],$$
 (2)

then

$$t_0 = START[i]. (3)$$

iii. If

$$t_0 \geq \text{STOP}[i],$$
 (4)

then

$$t_0 = STOP[i]. (5)$$

iv. If

$$j = 0, (6)$$

then

$$t_0 = START[i]. (7)$$

v.

$$t_1 = \text{TIME}_{\text{mtl}}[j] + (1 - \text{TIMEPIXR}_{\text{mtl}}) \times \text{TIMEDEL}_{\text{mtl}}.$$
 (8)

vi. If

$$t_1 < \text{START}[i],$$
 (9)

then

$$t_1 = \mathsf{START}[i]. \tag{10}$$

vii. If

$$t_1 \geq \text{STOP}[i],$$
 (11)

then

$$t_1 = STOP[i]. (12)$$

viii. If

$$j = N_{\text{mtl}} - 1, \tag{13}$$

then

$$t_1 = STOP[i]. (14)$$

ix.

$$\Delta t = t_1 - t_0. \tag{15}$$

x. For each temperature interval $k = 0, 1, ..., N_{\text{resp}} - 1$:

A. If

$$\text{FP_TEMP}_{\text{mtl}}[j] \geq \text{FP_TEMP}_{\min}[k] \text{ and}$$
 (16)

$$FP_TEMP_{mtl}[j] < FP_TEMP_{max}[k],$$
 (17)

then

$$\tau[k] = \tau[k] + \Delta t,\tag{18}$$

where $\tau[k]$ is the exposure time associated with temperature interval k.

- 6. The ARF A is computed using the entire set of good-time intervals.
- 7. The good-time intervals GTI[k] are obtained for each temperature interval k.
- 8. A WMAP is obtained for each temperature interval k.
- 9. The RMF R_k is obtained for each temperature interval k using $\mathtt{respfile}[k]$ and the appropriate WMAP.
- 10. The ARF A_k is obtained for each temperature interval k using the appropriate WMAP.
- 11. The temperature-weighted RMF is given by 1

$$R = \frac{\sum_{k=0}^{N_{\text{resp}}-1} \tau[k] A_k R_k}{A \sum_{k=0}^{N_{\text{resp}}-1} \tau[k]}.$$
 (19)

 $^{^{1}\}mathrm{The}$ source spectrum is assumed to be the same for each temperature interval.