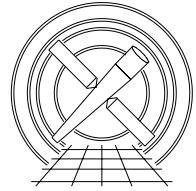




MIT Kavli Institute



Chandra X-Ray Center

MEMORANDUM

February 2, 2017

To: Jonathan McDowell, SDS Group Leader
From: Glenn E. Allen, SDS
Subject: Temperature-dependent RMF spec
Revision: 1.00
URL: <http://space.mit.edu/CXC/docs/docs.html#trmf>
File: /nfs/inconceivable/d0/sds/specs/mkacisrmf/trmf_spec.1.00.tex

1 Description

This spec is a high-level description of how a script could be used to create a temperature-dependent RMF.

2 Parameters

The parameters for the script are identical to the parameters for the tool `mkacisrmf`, except the script includes the two additional parameters:

1. `gtifile,f,a,"",,` “Input GTI file (e.g. the `evt2` file)”
2. `mtlfile,f,a,"",,` “Input mission timeline file”

3 Processing

1. The start and stop times for each one of the N_{gti} good-time intervals for the `CCD_ID` of interest are read from the appropriate GTI HDU of the `gtifile`. Hereafter these times for the i^{th} interval are referred to as `START` $[i]$ and `STOP` $[i]$, respectively.
2. The N_{mtl} sets of values for the time and focal-plane temperature are read from the `mtlfile`. Hereafter, the j^{th} time and temperature are referred to as `TIME` $_{\text{mtl}}[j]$ and `FP_TEMP` $_{\text{mtl}}[j]$, respectively. The header of the `mtlfile` is read to obtain the values of the keywords `TIMEDEL` and `TIMEPIXR`, which are hereafter referred to as `TIMEDEL` $_{\text{mtl}}$ and `TIMEPIXR` $_{\text{mtl}}$, respectively.
3. The CALDB is queried to obtain the N_{resp} response files associated with the time of the observation. There is one file for each temperature range calibrated. Hereafter, the k^{th} response file is referred to as `respfile` $[k]$.

4. The minimum and maximum focal-plane temperatures for each **respfile** are obtained from header keywords in the file. Hereafter these temperatures for the k^{th} file are referred to as, $\text{FP_TEMP}_{\min}[k]$ and $\text{FP_TEMP}_{\max}[k]$, respectively.
5. For each good-time interval $i = 0, 1, \dots, N_{\text{gti}} - 1$:

- (a) For each focal-plane temperature $j = 0, 1, \dots, N_{\text{mtl}} - 1$:

i.

$$t_0 = \text{TIME}_{\text{mtl}}[j] - \text{TIMEPIXR}_{\text{mtl}} \times \text{TIMEDEL}_{\text{mtl}}. \quad (1)$$

ii. If

$$t_0 < \text{START}[i], \quad (2)$$

then

$$t_0 = \text{START}[i]. \quad (3)$$

iii. If

$$t_0 \geq \text{STOP}[i], \quad (4)$$

then

$$t_0 = \text{STOP}[i]. \quad (5)$$

iv. If

$$j = 0, \quad (6)$$

then

$$t_0 = \text{START}[i]. \quad (7)$$

v.

$$t_1 = \text{TIME}_{\text{mtl}}[j] + (1 - \text{TIMEPIXR}_{\text{mtl}}) \times \text{TIMEDEL}_{\text{mtl}}. \quad (8)$$

vi. If

$$t_1 < \text{START}[i], \quad (9)$$

then

$$t_1 = \text{START}[i]. \quad (10)$$

vii. If

$$t_1 \geq \text{STOP}[i], \quad (11)$$

then

$$t_1 = \text{STOP}[i]. \quad (12)$$

viii. If

$$j = N_{\text{mtl}} - 1, \quad (13)$$

then

$$t_1 = \text{STOP}[i]. \quad (14)$$

ix.

$$\Delta t = t_1 - t_0. \quad (15)$$

x. For each temperature interval $k = 0, 1, \dots, N_{\text{resp}} - 1$:

A. If

$$\text{FP_TEMP}_{\text{mtl}}[j] \geq \text{FP_TEMP}_{\text{min}}[k] \text{ and} \quad (16)$$

$$\text{FP_TEMP}_{\text{mtl}}[j] < \text{FP_TEMP}_{\text{max}}[k], \quad (17)$$

then

$$\tau[k] = \tau[k] + \Delta t, \quad (18)$$

where $\tau[k]$ is the exposure time associated with temperature interval k .

6. The ARF A is computed using the entire set of good-time intervals.
7. The good-time intervals $\text{GTI}[k]$ are obtained for each temperature interval k .
8. A WMAP is obtained for each temperature interval k .
9. The RMF R_k is obtained for each temperature interval k using `respfile[k]` and the appropriate WMAP.
10. The ARF A_k is obtained for each temperature interval k using the appropriate WMAP.
11. The temperature-weighted RMF is given by¹

$$R = \frac{\sum_{k=0}^{N_{\text{resp}}-1} \tau[k] A_k R_k}{A \sum_{k=0}^{N_{\text{resp}}-1} \tau[k]}. \quad (19)$$

¹The source spectrum is assumed to be the same for each temperature interval.